




Riroriro: Simulating gravitational waves and calculating SNRs in Python

Wouter G. J. van Zeist¹, H  lo  se F. Stevance¹, and J. J. Eldridge¹

¹ Department of Physics, University of Auckland, New Zealand

DOI: [10.21105/joss.02968](https://doi.org/10.21105/joss.02968)

Software

- [Review](#) 
- [Repository](#) 
- [Archive](#) 

Editor: [Dan Foreman-Mackey](#) 

Reviewers:

- [@GregoryAshton](#)
- [@katiebreivik](#)

Submitted: 09 December 2020

Published: 15 February 2021

License

Authors of papers retain copyright and release the work under a Creative Commons Attribution 4.0 International License ([CC BY 4.0](#)).

Summary

Riroriro is a Python package to simulate the gravitational waveforms of binary mergers of black holes and/or neutron stars, and calculate several properties of these mergers and waveforms, specifically relating to their observability by gravitational wave detectors.

The gravitational waveform simulation of Riroriro is based upon the methods of [Buskirk & Babiuc-Hamilton \(2019\)](#), a paper which describes a computational implementation of an earlier theoretical gravitational waveform model by [Huerta et al. \(2017\)](#), using post-Newtonian expansions and an approximation called the implicit rotating source to simplify the Einstein field equations and simulate gravitational waves. Riroriro's calculation of signal-to-noise ratios (SNR) of gravitational wave events is based on the methods of [Barrett et al. \(2018\)](#), with the simpler gravitational wave model Findchirp ([Allen et al., 2012](#)) being used for comparison and calibration in these calculations.

Statement of Need

Gravitational waves have long been an area of research in astronomy, and particularly since the first observation of gravitational waves was announced in 2016 ([Abbott et al., 2016](#)) this area has seen a lot of research activity. Observations of gravitational waves from binary mergers can provide unique information about their progenitors and stellar populations, especially when combined with electromagnetic observations in the field called multi-messenger astronomy. A major factor in the successful detection and analysis of gravitational wave signals is the creation of simulations of such signals which observed data can be compared to. Because of this, multiple gravitational wave models have been created over the years. In particular, the gravitational wave observatories LIGO and Virgo have created their own models to use as templates in gravitational wave searches, with the main software for this being LALSuite ([LIGO Scientific Collaboration, 2018](#)). Various research groups have also created waveform models, with some examples of recent sophisticated waveform models being IMRPhenomXPHM ([Pratten et al., 2020](#)) and SEOBNRv4PHM ([Ossokine et al., 2020](#)).

We have not tested if the waveform model of Riroriro that is based on [Huerta et al. \(2017\)](#) and [Buskirk & Babiuc-Hamilton \(2019\)](#) is accurate enough to use for parameter estimation of detected gravitational wave transients as this was not within the scope of our project; it is likely that accurate parameter estimation requires careful modelling of the ringdown phase, especially for the most massive mergers. However, we use a level of accuracy adequate for our aim of modelling the detectability of gravitational wave transients predicted by stellar population syntheses. Furthermore, the code of Riroriro is structured and commented in such a way that each step in the process of the simulation is individually identifiable and modifiable by users. Users could also substitute in functions from other algorithms or even

40 use the detectability modules on waveforms from other sources, as long as the user puts these
41 in Riroriro's format.

42 Riroriro combines areas covered by previous models (such as gravitational wave simulation,
43 SNR calculation, horizon distance calculation) into a single package with broader scope and
44 versatility in Python, a programming language that is ubiquitous in astronomy. Aside from
45 being a research tool, Riroriro is also designed to be easy to use and modify, and it can
46 also be used as an educational tool for students learning about gravitational waves.

47 Features

48 Features of Riroriro include:

- 49 ▪ Simulating the gravitational waveform signal from a binary merger of two black holes,
50 two neutron stars or a black hole and a neutron star and outputting the data of this
51 signal in terms of frequency and strain amplitude.
- 52 ▪ Using a gravitational wave output and given a detector noise spectrum (such spectra
53 are made publicly available by LIGO), calculating the signal-to-noise ratio (SNR) of the
54 signal at a given distance assuming optimal alignment.
- 55 ▪ Calculating the horizon distance (maximum distance at which an event could be ob-
56 served) for a gravitational wave model and a given detector.
- 57 ▪ Given the optimal-alignment SNR of an event, evaluating its detectability, the probability
58 that the event would be detected with a SNR above the commonly used threshold of
59 8, if the alignment would be arbitrary. These results could then be combined with
60 population synthesis calculations to estimate how many of the predicted mergers would
61 be detected.

62 In addition, we have created Jupyter Notebook tutorials to help users get started with Riror
63 iro ([see tutorials](#)).

64 Research

65 Riroriro has been used for research in conjunction with BPASS, a suite of computer programs
66 that simulates the evolution of a population of binary and single-star systems from a wide
67 range of initial conditions and predicts their electromagnetic spectral emission ([Eldridge et
68 al., 2017](#); [Stanway & Eldridge, 2018](#)). There is also a Python interface for BPASS called Hoki
69 ([Stevance et al., 2020](#)). This research took rates of formation of merging systems from BPASS
70 and then evaluated the detectability of the gravitational wave signals from those systems using
71 Riroriro. This was done to obtain predictions of the rates at which gravitational waves of
72 different types would be expected to be observed, which can then be directly compared to those
73 events found by the LIGO/Virgo gravitational wave observatories ([Ghodla et al., forthcoming](#)).

74 Acknowledgments

75 HFS and JJE acknowledge support from the University of Auckland and also the Royal Society
76 of New Zealand Te Apārangi under the Marsden Fund.

References

- Abbott, B. P., Abbott, R., Abbott, T. D., Abernathy, M. R., Acernese, F., Ackley, K., Adams, C., Adams, T., Addesso, P., Adhikari, R. X., Adya, V. B., Affeldt, C., Agathos, M., Agatsuma, K., Aggarwal, N., Aguiar, O. D., Aiello, L., Ain, A., Ajith, P., ... Virgo Collaboration. (2016). Observation of gravitational waves from a binary black hole merger. *Phys. Rev. Lett.*, 116(6), 061102. <https://doi.org/10.1103/PhysRevLett.116.061102>
- Allen, B., Anderson, W. G., Brady, P. R., Brown, D. A., & Creighton, J. D. E. (2012). FINDCHIRP: An algorithm for detection of gravitational waves from inspiraling compact binaries. *Phys. Rev. D*, 85(12), 122006. <https://doi.org/10.1103/PhysRevD.85.122006>
- Barrett, J. W., Gaebel, S. M., Neijssel, C. J., Vigna-Gómez, A., Stevenson, S., Berry, C. P. L., Farr, W. M., & Mandel, I. (2018). Accuracy of inference on the physics of binary evolution from gravitational-wave observations. *MNRAS*, 477(4), 4685–4695. <https://doi.org/10.1093/mnras/sty908>
- Buskirk, D., & Babiuc-Hamilton, M. C. (2019). A complete analytic gravitational wave model for undergraduates. *European Journal of Physics*, 40(2), 025603. <https://doi.org/10.1088/1361-6404/aaf81e>
- Eldridge, J. J., Stanway, E. R., Xiao, L., McClelland, L. A. S., Taylor, G., Ng, M., Greis, S. M. L., & Bray, J. C. (2017). Binary Population and Spectral Synthesis Version 2.1: Construction, observational verification, and new results. *Publ. Astron. Soc. Aust.*, 34, e058. <https://doi.org/10.1017/pasa.2017.51>
- Ghodla, S., van Zeist, W. G. J., Eldridge, J. J., Stevance, H. F., & Stanway, E. R. (forthcoming). *Predicting the LIGO/VIRGO O3a GW transient mass distributions with BPASS*. Forthcoming.
- Huerta, E. A., Kumar, P., Agarwal, B., George, D., Schive, H.-Y., Pfeiffer, H. P., Haas, R., Ren, W., Chu, T., Boyle, M., Hemberger, D. A., Kidder, L. E., Scheel, M. A., & Szilágyi, B. (2017). Complete waveform model for compact binaries on eccentric orbits. *Phys. Rev. D*, 95(2), 024038. <https://doi.org/10.1103/PhysRevD.95.024038>
- LIGO Scientific Collaboration. (2018). *LIGO Algorithm Library - LALSuite*. Free software (GPL). <https://doi.org/10.7935/GT1W-FZ16>
- Ossokine, S., Buonanno, A., Marsat, S., Cotesta, R., Babak, S., Dietrich, T., Haas, R., Hinder, I., Pfeiffer, H. P., Pürrer, M., Woodford, C. J., Boyle, M., Kidder, L. E., Scheel, M. A., & Szilágyi, B. (2020). Multipolar effective-one-body waveforms for precessing binary black holes: Construction and validation. *Phys. Rev. D*, 102(4), 044055. <https://doi.org/10.1103/PhysRevD.102.044055>
- Pratten, G., García-Quirós, C., Colleoni, M., Ramos-Buades, A., Estellés, H., Mateu-Lucena, M., Jaume, R., Haney, M., Keitel, D., Thompson, J. E., & Husa, S. (2020). Let's twist again: Computationally efficient models for the dominant and sub-dominant harmonic modes of precessing binary black holes. *arXiv e-Prints*, arXiv:2004.06503. <http://arxiv.org/abs/2004.06503>
- Stanway, E. R., & Eldridge, J. J. (2018). Re-evaluating old stellar populations. *MNRAS*, 479(1), 75–93. <https://doi.org/10.1093/mnras/sty1353>
- Stevance, H. F., Eldridge, J. J., & Stanway, E. R. (2020). Hoki: Making BPASS accessible through Python. *The Journal of Open Source Software*, 5(45), 1987. <https://doi.org/10.21105/joss.01987>